

Paper Number: 3996

Classification, mineralogical and geochemical variations in pegmatites of the Cape Cross –Uis pegmatite belt, Namibia.

Fuchsloch, W.C.¹, Nex, P.A.M. ¹, and Kinnaird, J.A.¹



1 School of Geosciences, University of the Witwatersrand, South Africa. wfuchsloch@gmail.com

The Pan African Damara Orogen, Namibia, is host to a variety of mineral deposits which include pegmatite hosted lithium, tin, niobium-tantalum, uranium, and semi-precious stones (Miller, 1983[3]; Diehl, 1993[2]; Ashworth, 2013[1]). This paper is focussed on examining syn- and post-orogenic pegmatites (~490 Ma; mineralised and barren) of the Cape Cross–Uis pegmatite belt, to understand and give insight into mineralisation controls and pegmatite evolution. The Cape Cross–Uis pegmatite belt is one of ten narrow regional pegmatite belts, trending parallel to the orogen and lies within the Northern and South Kaoko Zones of the tectonostratigraphic zones of the Damara Orogen (Miller, 1983[3]). Within the belt, 3 major pegmatite swarms have been identified (Diehl, 1993[2]); the Strathmore, Karlowa and Uis Swarms. These swarms include cassiterite- and tantalite-bearing, unzoned, Uis-type tin pegmatites, and zoned lithium-beryllium pegmatites, barren pegmatites and weakly zoned tourmaline-bearing, intragranitic pegmatites (groups 1, 2, 3 and 4 respectively)

Field mapping evidence shows that pegmatites intrude with a NNE trend, in line with the regional structural fabric, along fold axes and bedding plane discontinuities. Pegmatites are variable in shape and size but are generally elongate and stock-work like in form and range from 10 – 300 m in length. Thickness also varies from about 1–70 m. Whole-rock major and trace element data has been acquired for pegmatites and granite and metasediment host rocks. Fine grained border zones of pegmatites and later stage greisen and albatized zones/pods were analysed to attain a possible chemical change between initial crystallisation and late stage crystallisation and fluid interaction. Granites and group 4 pegmatites show similar geochemical characteristics, such as elevated TiO₂, Ba, Y, Th, U, and total REE values compared to negligible amounts (< 10 ppm) in other pegmatite groups. Metasediment hosted pegmatites (group 1 – 3) show enrichment compared to granites and group 1 pegmatites in most other trace element values such as K₂O, Ta, Sn, Cs, Rb, Nb and peraluminosity. Geochemical evidence suggests that metasediment hosted pegmatites do not share a parent daughter relationship with either Salem-type or Red/Grey granites. Metasediment host rock interaction is evident within the pegmatite belt. Microliths of tourmaline and a bleaching of host rocks is a result of this interaction and a large increase in the Sn, Ta, and Nb compositions of the host rock. Whether or not this can be attributed to pegmatite fluids remobilising Ta, Nb and Sn and concentrating them in greisenised areas (assimilation), is debatable. The reverse might be more plausible, where late stage fractionated pegmatitic fluids rich in Ta, Nb, and Sn, leached into the host rocks, causing tourmalinisation and enriched ore values. Processes such as fractionation can be traced by whole-rock geochemistry as a definitive contributor to mineralisation and pegmatite evolution. Tantalite and cassiterite microprobe analyses are near completion.

[1] Ashworth, L. (2013). Mineralised pegmatites of the Damara Belt, Namibia. Fluid inclusion and geochemical characteristics with implications of post tectonic mineralisation. PhD thesis (unpublished), University of the Witwatersrand, Johannesburg.

[2] Diehl, B.T.M. (1993). Rare metal pegmatites of the Cape Cross – Uis pegmatite belt, Namibia. *Geology*,

mineralisation, rubidium strontium Characteristics and petrogenesis. *Journal of African earth sciences*, **17**, 167–181.

[3] Miller, R McG. 1983: The Pan-African Damara Orogen of South West Africa/Namibia. In: R McG Miller (Ed.): The Evolution of the Damara Orogen of South West Africa/Namibia. *Geological Society of South Africa Special Publication* **11**, 431-515.

